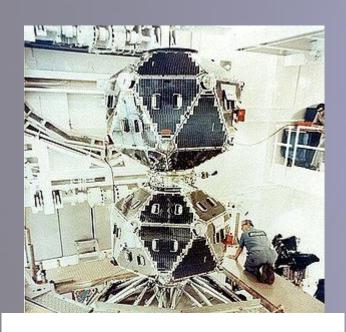
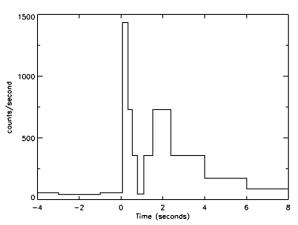


Gamma Ray Astronomy Beginning





started as a small budget research program in 1959

monitoring compliance with the 1963 Partial Test Ban Treaty by the Soviet Union, and other nuclearcapable states

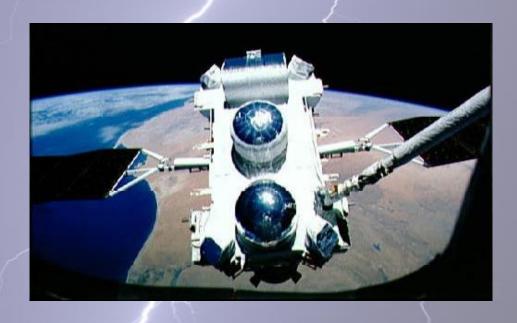
12 external X-ray detectors and 18 internal neutron and gamma-ray detectors

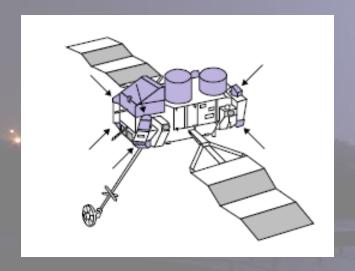
on July 2, 1967, at 14:19 UTC, the Vela 4 and Vela 3 satellites detected a flash of gamma radiation...

Compton Gamma Ray Observatory (CGRO)

Burst And Transient Source Experiment (BATSE, 1991-2000):

eight scintillator panels
each with effective area 2000 cm²
energies from 25 keV to above 1MeV
mounted on the corners of CGRO (spatial information)

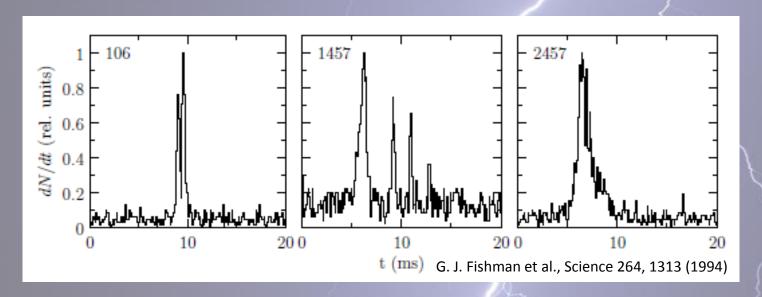




Sometimes bursts of radiation were observed only by detectors facing the Earth.

They were much shorter than the typical gammaray burst

Compton Gamma Ray Observatory (CGRO)



The y axis represents count rate of 25-1000 keV photons

BATSE data shows TGFs as short bursts of up to 1000 photons with energies ranging from 25 keV up to above 1MeV.

About 70 events were detected during the entire mission

Ramaty High Energy Solar Spectroscopic Imager



Small Explorer (SMEX)

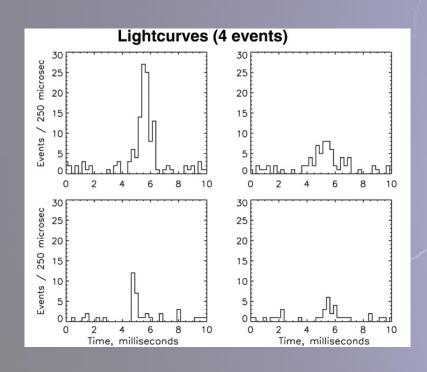
No shielding (good for TGF, bad for solar flares)

Each photon is counted – time and energy

TGFs recognized through time characteristic – duration less than 10 ms is much shorter than for typical GRB

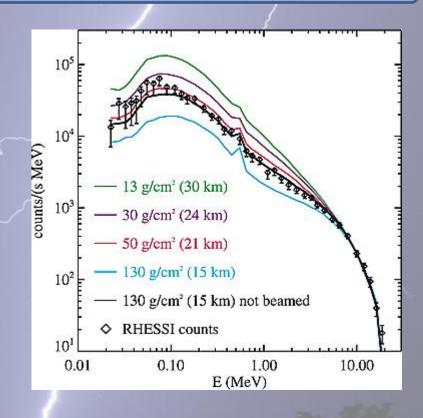
RHESSI does not give spatial information concernig TGFs

Ramaty High Energy Solar Spectroscopic Imager



Worse sensitivity (effective area 200 cm²)
Better time resolution

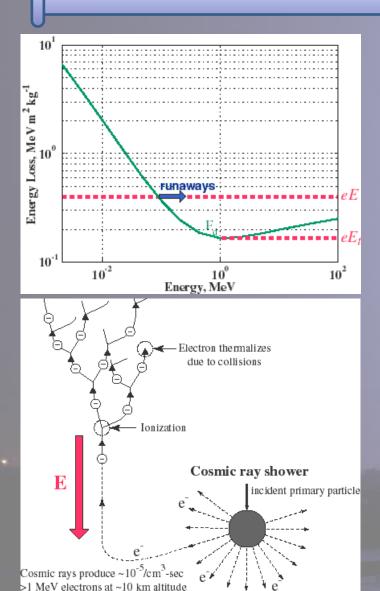
Over 800 events (years 2002 - 2008)



Accumulated spectrum of 85 RHESSI TGFs, with theoretical relativistic-runaway spectra

The only difference between the models is the altitude of the average TGF

Runaway Breakdown Mechanism



Gurevich et al., Phys. Letters A 165, 463 (1992)

Energy loss due to collisions decrease with increasing energy

Electrons with sufficient initial energy can be accelerated by electric fields

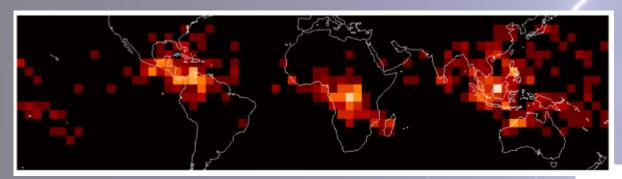
Accelerated electrons collide with atmospheric molecules, ionize them and produce more relativistic electrons.

Most newly produced electrons thermalize because of collisions (gamma ray emission), but some accelerate and contribute to the avalanche

Spark: cosmic ray particle

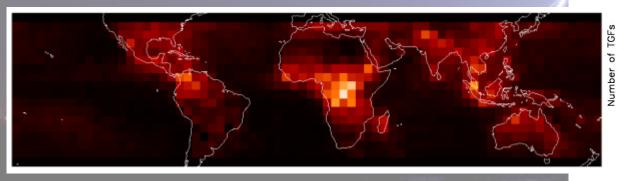
RHESSI TGFs and Lightnins

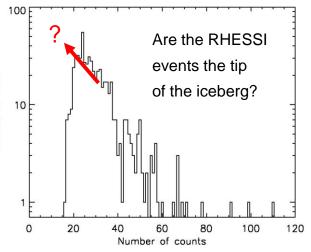
RHESSI TGF Positions



RHESSI TGFs for which lightning data is available show coincident lightning activity within several milliseconds for 76% of TGFs

Visible Lightning Positions





Ramaty High Energy Solar Spectroscopic Imager



What do You know about lightnings?

Blinding

Hit the ground

Short duration

Relatively simple shape

Additional effects: thunder

oh, Yeah

are You sure?

well, but...

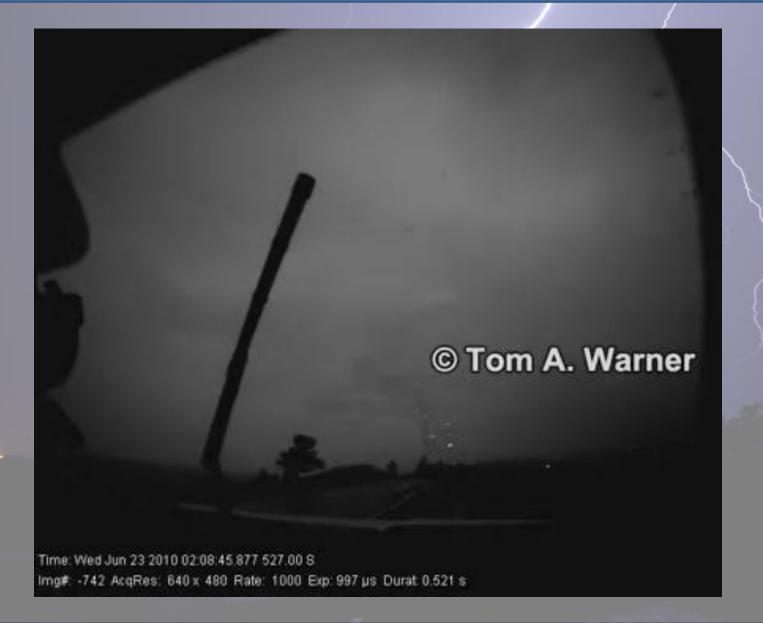
definately not

not only





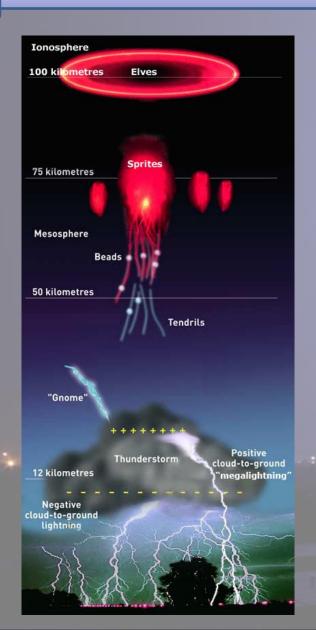
What do You know about lightnings?

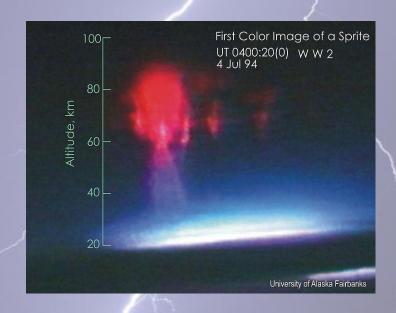


What do You know about lightnings?



Sprites, Elves, Gnomes...





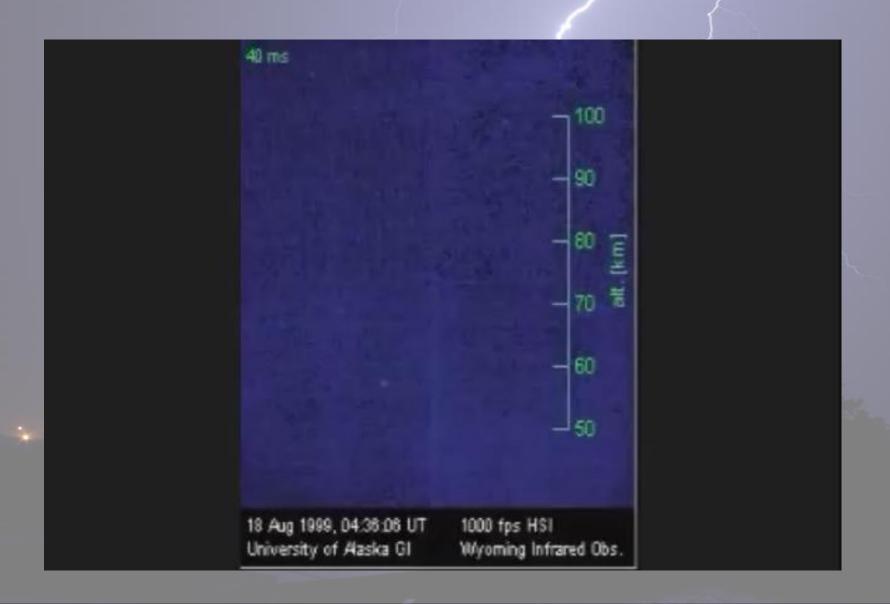
first photographed on July 6, 1989

observed above thunderstorms

extremely short

observed few miliseconds after lightning

Sprites



Sprites

Sprites



Consistent with running breakdown mechanism

The source of TGFs?

Electron thermalizes due to collisions

Cosmic ray shower

>1 MeV electrons at ~10 km altitude

incident primary particle

New Instrumentation



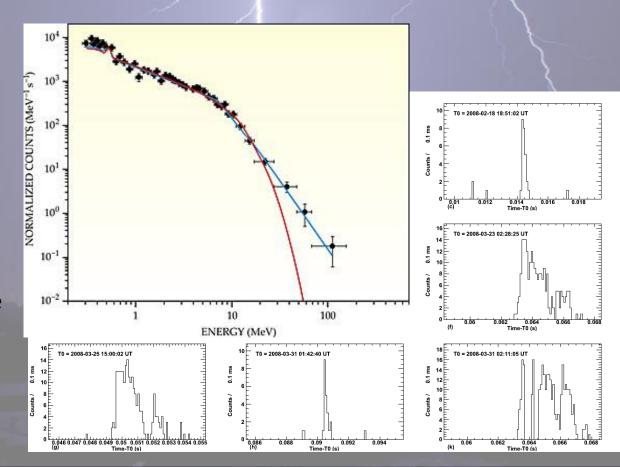
detects ~10 TGFs/month

AGILE and RHESSI TGF samples are consistent concerning longitude, local time distribution and spectral shape

cumulative spectrum with significant detection above 40 MeV

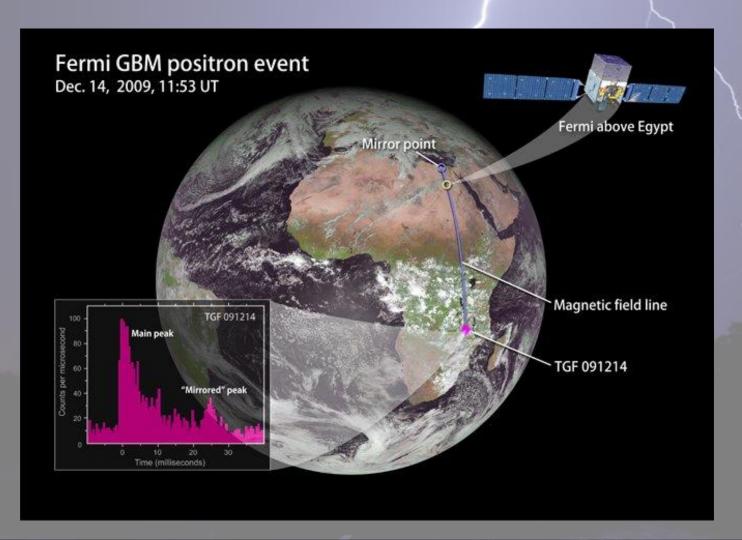
AGILE (Astro-rivelatore Gamma a Immagini LEggero)

combines gamma-ray imager (30 MeV- 30 GeV) with a hard X-ray imager (18-60 keV) with large FOVs (1-2.5 sr) and optimal angular resolution



New Instrumentation

Fermi Gamma Ray Space Telescope and new discovery



Conclusions We live in interesting times...